

# Backup slides

# Coherence production conditions

Coherence production conditions:

$$|\Delta E| \ll \sigma_E, \quad |\Delta p| \ll \sigma_p.$$

On the other hand:

$$\Delta E \simeq v_g \Delta p + \frac{\Delta m^2}{2E}.$$

Constraint  $|\Delta E| \ll \sigma_E \Rightarrow$

$$\left| \frac{v_g \Delta p}{\sigma_E} + \frac{\Delta m^2}{2E \sigma_E} \right| \ll 1. \quad (*)$$

(a) The two terms in  $\Delta E$  do not approximately cancel each other.  $\Rightarrow$

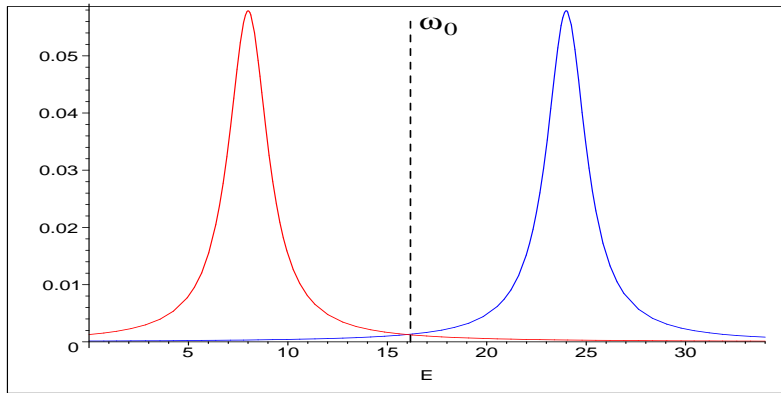
$v_g |\Delta p| \ll \sigma_E \leq \sigma_p$ , i.e. for relativistic neutrinos  $|\Delta p| \ll \sigma_p$  follows from  $|\Delta E| \ll \sigma_E$ .

(b1) There is a strong cancellation, but both terms on the l.h.s. of (\*) are small – see case (a).

(b2) Strong cancellation, but both terms on the l.h.s. of (\*) are  $\gtrsim 1$ : momentum condition is independent. But: the only known case – Mössbauer neutrinos.

# Mössbauer effect

Conventional Mössbauer effect – Res. absorption of  $\gamma$  quanta:



Nuclear exc. energy:  $\omega_0$ .

Recoil energy:  $R = \frac{\omega_0^2}{2M}$

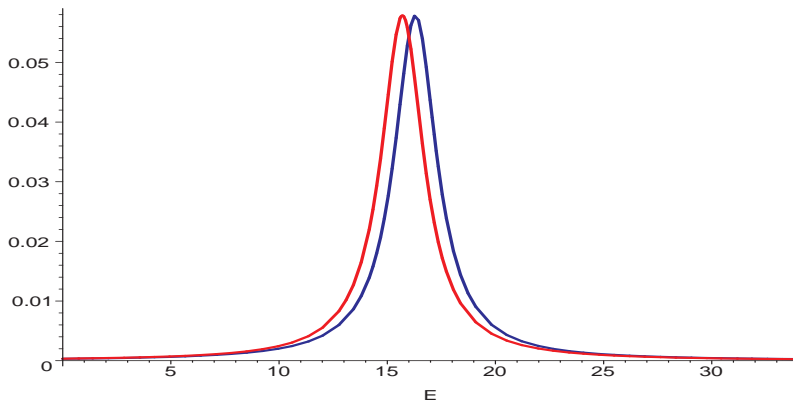
$$E_e = \omega_0 - \frac{\omega_0^2}{2M}$$

$$E_a = \omega_0 + \frac{\omega_0^2}{2M}$$

Recoilless emission and absorption (Mössb. eff.):

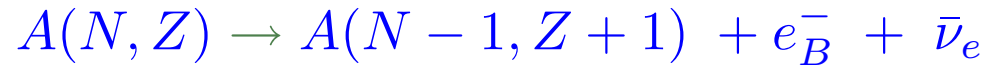
$$E_e \simeq E_a \simeq \omega_0$$

Strong enhancement of absorption

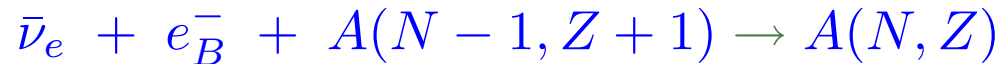


# Mössbauer effect with neutrinos?

Beta decay with 2 - body final state:



Inverse process:



If the nuclei are embedded in solid state lattice, recoilless emission and absorption in principle possible.

Possibility of Mössbauer effect with neutrinos:

Visscher, 1959; Kells & Schiffer, 1983; Raghavan, 2005, 2006

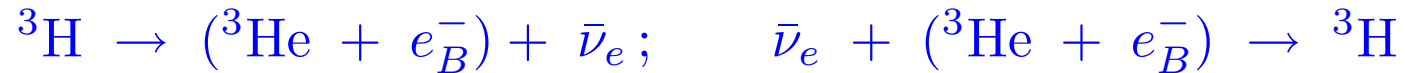
Relevant processes considered:

Bahcall, 1961 – bound state  $\beta$  decay;

Mikaelyan, Tsinoev & Borovoi, 1967 – inverse process  
(stimulated K-electron capture)

# Mössbauer effect with neutrinos?

Mössbauer effect with neutrinos on  ${}^3\text{H} - {}^3\text{He}$  system:



Energy release:  $Q = 18.6 \text{ keV}$ . Mean lifetime of  ${}^3\text{H}$  is 17.8 yr  $\Rightarrow$

Nat. linewidth  $\Gamma_{{}^3\text{H}} = 1.17 \times 10^{-24} \text{ eV}$  – extremely small:  $\Delta E/E \sim 10^{-28}$  !

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Various (homogeneous and inhomogeneous) broadening effects exist. By suppressing them probably an effective linewidth  $\Gamma_{\text{eff}} \sim 10^{-11} \text{ eV}$  can be achieved (W. Potzel)  $\Rightarrow \Delta E/E \sim 10^{-15}$  – still very small.

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Number of  ${}^3\text{H}$  atoms produced in the target can be counted by detecting their decay or using mass spectroscopy.

Very serious technical difficulties exist, but apparently realization of a Mössbauer experiment with neutrinos is not impossible (Raghavan, Potzel).

If realized: for  $\Gamma \sim 10^{-11} \text{ eV}$ ,  $\sigma \sim 10^{-33} \text{ cm}^2$  !

# Mössbauer effect with neutrinos?

If a Mössbauer neutrino experiment is realized  $\Rightarrow$  a unique source of extremely monochromatic low energy neutrinos. Would open up possibilities

- to detect for the first time keV neutrinos
- to detect neutrinos with g or 100 g scale (rather than t or kt scale) detectors
- to observe gravitational redshift of neutrinos
- to study neutrino oscillations at distances  $\sim 10$  m rather than km or hundreds/thousands of km
- to search for the effects of yet unmeasured mixing angle  $\theta_{13}$  and possibly measure it
- to discriminate between the normal and inverted neutrino mass hierarchies without using matter effects
- to study possible oscillations into sterile neutrino states



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Arguments in the literature (Bilenky et al.):

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Is that true?

# Will Mössbauer neutrinos oscillate?

Neutrino oscillations require some intrinsic uncertainty of energy and momentum of the emitted and detected neutrino states !

If  $E$  and  $p$  were known precisely, from  $E^2 = p^2 + m_i^2$  one would determine which mass eigenstate has been emitted  $\Rightarrow$  neutrinos of different mass would not be emitted coherently.

For Mössbauer effect with neutrinos in  ${}^3\text{H} - {}^3\text{He}$  system:

$$\diamond \frac{\Delta m^2}{2E} = \frac{2.5 \times 10^{-3} \text{ eV}^2}{2 \cdot 18.6 \text{ keV}} \simeq 6.7 \cdot 10^{-8} \text{ eV} \gg \Gamma \sim 10^{-11} \text{ eV}!$$

Can neutrinos of different mass be accommodated within such a small energy uncertainty?

Will neutrinos with such small energy uncertainty oscillate ?

# Two “standard” approaches to $\nu$ oscillations

The oscillation phase:

$$\phi = p_\mu x^\mu = E \cdot t - p \cdot x \quad \Rightarrow$$

$$\Delta\phi = \Delta E \cdot t - \Delta p \cdot L$$

I. Same momentum approach ( $\Delta p = 0$ ). The oscillation phase

$$\Delta\phi = \Delta E \cdot t - \Delta p \cdot L \quad \Rightarrow \quad \Delta E \cdot t$$

– evolution in time; needs to use  $L \simeq t$ .

II. Same energy approach ( $\Delta E = 0$ ):

$$\Delta\phi = -\Delta p \cdot L$$

– evolution in space.

# Will Mössbauer neutrinos oscillate?

– Same momentum approach (evolution in time): no.

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Our point of view: in general, there is no reason to believe that  $\nu_i$  have either same energy or same momentum. No need to perform Mössbauer  $\nu$  experiment to decide which approach is correct – it is sufficient to carefully examine the validity of the approximations used.

# How about Mössbauer neutrinos?

Very small effective linewidth  $\Gamma \Rightarrow$  small energy uncertainty of the emitted neutrino state. Can different neutrino mass eigenstates be emitted coherently?

$$\sigma_{m^2} = [(2E\sigma_E)^2 + (2p\sigma_p)^2]^{1/2}$$

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$$\Rightarrow \sigma_p \sim 10 \text{ keV}, \quad \text{i.e.} \quad \sigma_{m^2}^2 \simeq 2p\sigma_p \sim 4 \times 10^8 \text{ eV}^2 \gg \Delta m^2$$

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$\Rightarrow$  **Oscillations must occur !**

# QFT calculation

Inhomogeneous line broadening: Calculate the probability of the overall process for zero linewidths and then average the result over the energy distribution of  ${}^3\text{H}$  and  ${}^3\text{He}$  nuclei in the source and detector.

Homogeneous line broadening: modify the amplitude of the process and apply a proper averaging procedure to take into account the stochastic nature of the processes leading to homog. broadening.  $\Rightarrow$  Results in both cases are formally very similar. Mössbauer res. condition:

$$|E_S - E_D| \ll \gamma_S + \gamma_D$$

If it is satisfied  $\Rightarrow$  neutrino detection cross section enhanced by a factor

$$\sim (\alpha Z m_e)^3 / [p_e E_e (\gamma_S + \gamma_D)] \sim 10^{12}$$

compared to non-resonance  $\sigma(\bar{\nu}_e + A \rightarrow A' + e^+)$  for neutrinos of same energy (assuming recoil-free fraction  $\sim 1$ ).



# QFT calculation – contd.

The amplitude for zero linewidths:

$$\begin{aligned}
 i\mathcal{A} = & \int d^3x_1 dt_1 \int d^3x_2 dt_2 \Psi_{He,S}^*(\vec{x}_1) e^{+iE_{He,S} t_1} \Psi_{H,S}(\vec{x}_1) e^{-iE_{H,S} t_1} \\
 & \cdot \Psi_{H,D}^*(\vec{x}_2) e^{+iE_{H,D} t_2} \Psi_{He,S}(\vec{x}_2) e^{-iE_{He,D} t_2} \\
 & \cdot \sum_j \mathcal{M}_S^\mu \mathcal{M}_D^{\nu*} |U_{ej}|^2 \int \frac{d^4p}{(2\pi)^4} e^{-ip_0(t_2-t_1) + i\vec{p}(\vec{x}_2-\vec{x}_1)} \\
 & \cdot \bar{u}_{e,S} \gamma_\mu (1 - \gamma_5) \frac{i(\not{p} + m_j)}{p_0^2 - \vec{p}^2 - m_j^2 + i\epsilon} \gamma_\nu (1 - \gamma_5) u_{e,D}
 \end{aligned}$$

Here

$$\mathcal{M}_{S,D}^\mu = \frac{G_F \cos \theta_c}{\sqrt{2}} \psi_e(R) \bar{u}_{He} (M_V \delta_0^\mu - g_A M_A \sigma_i \delta_i^\mu / \sqrt{3}) u_H \kappa_{S,D}^{1/2}$$

# QFT calculation – contd.

The overall process rate:

$$\Gamma = \frac{\Gamma_0 B_0}{4\pi L^2} Y_S Y_D \int_0^\infty dE_{H,S} dE_{He,S} dE_{He,D} dE_{H,D}$$
$$\cdot \delta(E_S - E_D) \rho_{H,S}(E_{H,S}) \rho_{He,D}(E_{He,D}) \rho_{He,S}(E_{He,S}) \rho_{H,D}(E_{H,D})$$
$$\cdot \sum_{j,k} |U_{ej}|^2 |U_{ek}|^2 \exp \left[ -\frac{2E_S^2 - m_j^2 - m_k^2}{2\sigma_p^2} \right] e^{i(\sqrt{E_S^2 - m_j^2} - \sqrt{E_S^2 - m_k^2})L}$$

$\sigma_p$  – effective momentum uncertainty of the emission/absorption processes:

$$\frac{1}{\sigma_p^2} = \frac{1}{m_H \omega_{H,S} + m_{He} \omega_{He,S}} + \frac{1}{m_H \omega_{H,D} + m_{He} \omega_{He,D}},$$

An analogue of the Debye - Waller (Lamb - Mössbauer) factor:

$$\diamond \exp[-(2E_S^2 - m_j^2 - m_k^2)/2\sigma_p^2] = \exp[-(p_j^2 + p_k^2)/2\sigma_p^2]$$

# QFT calculation – contd.

For Lorentzian energy distributions of external particles:

$$\rho_{A,B}(E_{A,B}) = \frac{\gamma_{A,B}/2\pi}{(E_{A,B} - E_{A,B,0})^2 + \gamma_{A,B}^2/4}$$

$$(A = \{H, He\}, B = \{S, D\}, E_{A,B,0} = m_A + \frac{1}{2}\omega_{A,B}) \Rightarrow$$

$$\Gamma \simeq \frac{\Gamma_0 B_0}{4\pi L^2} Y_S Y_D \sum_{j,k} |U_{ej}|^2 |U_{ek}|^2 \exp\left[-\frac{(p_{jk}^{\min})^2}{\sigma_p^2}\right] \exp\left[-\frac{|\Delta m_{jk}^2|}{2\sigma_p^2}\right]$$

$$\cdot \frac{1}{2} \left( e^{-L/L_{jk,S}^{\text{coh}}} + e^{-L/L_{jk,D}^{\text{coh}}} \right) \exp\left[-i \frac{\Delta m_{jk}^2}{2\bar{E}} L\right] \frac{(\gamma_S + \gamma_D)/2\pi}{(E_{S,0} - E_{D,0})^2 + \frac{(\gamma_S + \gamma_D)^2}{4}}$$

$L_{jk,B}^{\text{coh}}$  – coherence lengths:

$$L_{jk,B}^{\text{coh}} = \frac{4\bar{E}^2}{\gamma_B |\Delta m_{jk}^2|} = \frac{\sigma_x}{\Delta v_g}, \quad \sigma_x = \frac{2}{\gamma_B} \quad (B = S, D)$$

# QFT calculation – contd.

Generalized Lamb – Mössbauer (Debye – Waller) factor

$$\exp \left[ -\frac{p_j^2 + p_k^2}{2\sigma_p^2} \right] = \exp \left[ -\frac{(p_{jk}^{\min})^2}{\sigma_p^2} \right] \exp \left[ -\frac{|\Delta m_{jk}^2|}{2\sigma_p^2} \right]$$

First factor  $\Rightarrow$  suppression of emission and absorption, i.e. a generalized Lamb-Mössbauer factor, second factor  $\Rightarrow$  suppression of oscillations.

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$|\Delta m_{jk}^2| \lesssim 2\sigma_p^2 \Rightarrow$  localization condition: Spatial localization  $\sigma_x \sim 1/\sigma_p$ .

Oscillations would be suppressed only if  $|\Delta m_{jk}^2| \gtrsim 2\sigma_p^2$ .

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In reality:  $|\Delta m_{jk}^2|_{\max} \simeq 2.5 \cdot 10^{-3} \text{ eV}^2$ ;  $\sigma_p^2 \sim (10 \text{ keV})^2 \Rightarrow$

oscillations will not be suppressed.

# QFT calculation – contd.

For realistic values of parameters – just the expected result: the rate of no-oscillation production-detection process times the standard oscillation probability (probability of  $\bar{\nu}_e$  survival). Decoherence and delocalization can be neglected.

## Conclusion:

If a Mössbauer neutrino experiment is realized – recoillessly emitted and absorbed neutrinos will oscillate.





# Longitudinal vs. transversal w.p. dispersion

Spreading of the wave packets: consequence of the fact that there is a spread of momenta inside of the wave packets and of the  $p$ -dependence of the group velocity.

$$v_{spr}^i \simeq \frac{\partial v_i}{\partial p^j} \sigma_p^j = \frac{1}{E} (\delta_{ij} - v_i v_j) \sigma_p^j = \frac{1}{E} [\sigma_p^i - v_i (\vec{v} \vec{\sigma}_p)]$$

This gives

$$v_{spr.}^\perp = \frac{\sigma_p}{E}, \quad v_{spr.}^\parallel = \frac{\sigma_p}{E} (1 - v^2) = \frac{\sigma_p}{E} \frac{m^2}{E^2}$$

$$t_{transv} \sim E/\sigma_p^2, \quad t_{long.} \sim E^3/\sigma_p^2 m^2.$$