
Physics at e^+e^- colliders: present and future

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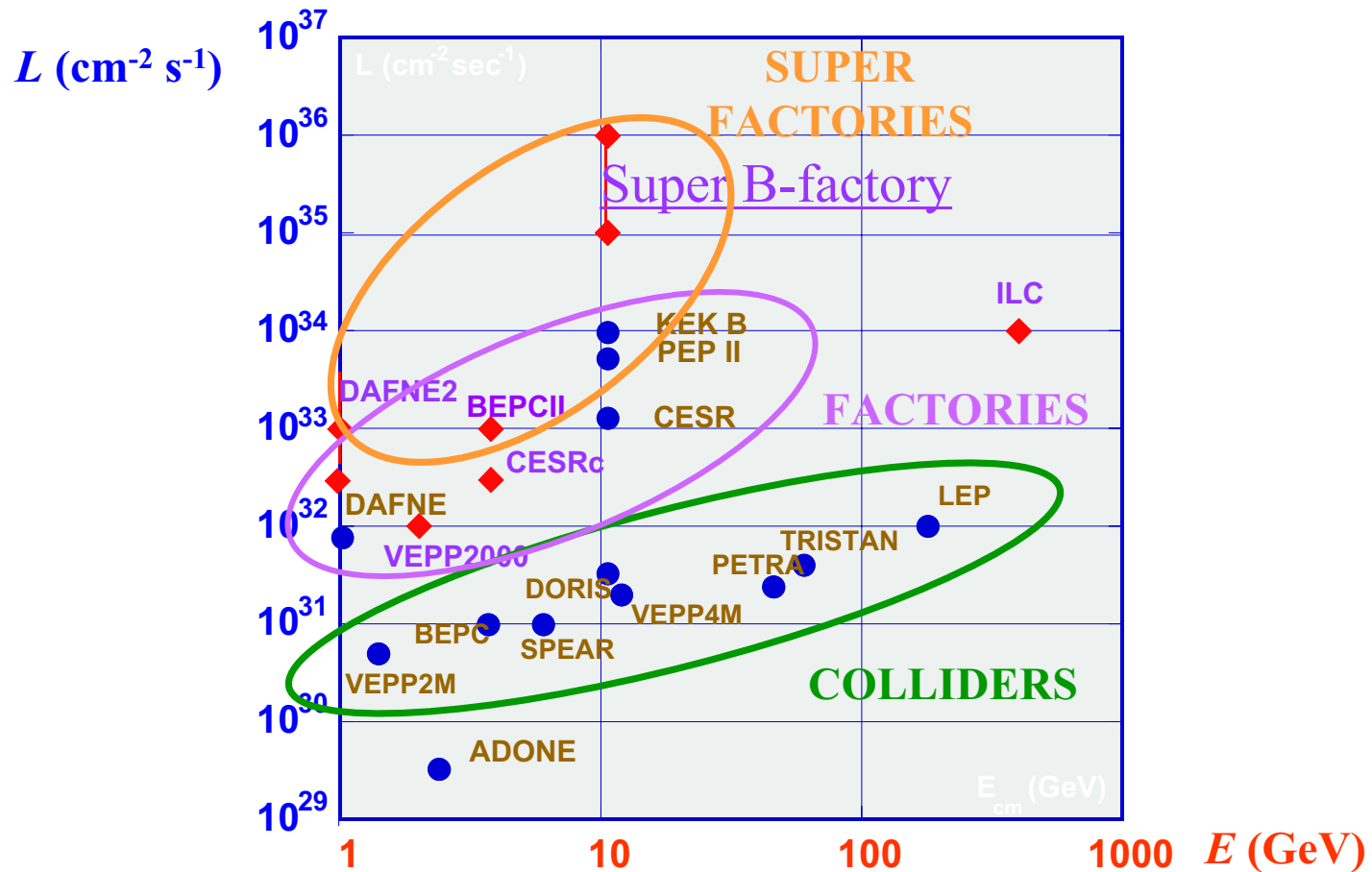
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Traditional purposes for e^+e^- colliders

- ✓ Vector mesons $J^P = 1^-$ ($\rho, \omega \dots Z^0$) and their decays
- ✓ $e^+e^- \rightarrow \text{hadrons}$
- ✓ Two-photon hadron production
- ✓ Studying of $c\bar{c}$ and $b\bar{b}$ families
- ✓ Cabibbo-Kobayashi-Maskawa (CKM) mixing

High precision experiments of last years added new physical goals

- Search of oscillations in heavy quarks systems
- Search of CP-violation with c- and b-quarks.
- Search of New Physics effects. Most sound example is anomalous magnetic moment of muon.



Main parameter of e^+e^- collider is luminosity L

$$N = L \times \sigma$$

$$\vec{\mu} = g_\mu \frac{e\hbar}{2m_\mu c} \vec{s} \quad (1)$$

From Dirac equation $g_\mu = 2$ but radiative corrections change it

$$g_\mu = 2 \cdot \left(1 + \frac{\alpha}{2\pi} + \dots\right) \quad \text{Schwinger, 1948} \quad (2)$$

Anomalous magnetic moment of muon:

$$a_\mu = (g_\mu - 2)/2$$

It is measured with high precision

$$a_{\mu^+}^{exp} = 116592030(80) \times 10^{-11} \quad 2002 \quad E821 \text{ at BNL}$$

$$a_{\mu^-}^{exp} = 116592140(80) \times 10^{-11} \quad 2004$$

Average:

$$a_\mu^{exp} = 116592080(60) \times 10^{-11}$$

General form of lepton-photon vertex:

$$\bar{u}(p')\Gamma_\mu u(p) = \bar{u}(p') \left[\gamma_\mu F_1(q^2) + \frac{i}{2m} \sigma_{\mu\nu} q^\nu F_2(q^2) \right] u(p)$$

$$F_1(0) = 1, \quad F_2(0) = a$$

The Standard Model prediction for a_μ can be represented as a sum

$$a_\mu^{SM} = a_\mu^{QED} + a_\mu^{had} + a_\mu^{EW}$$

QED contribution contains only leptons and photons

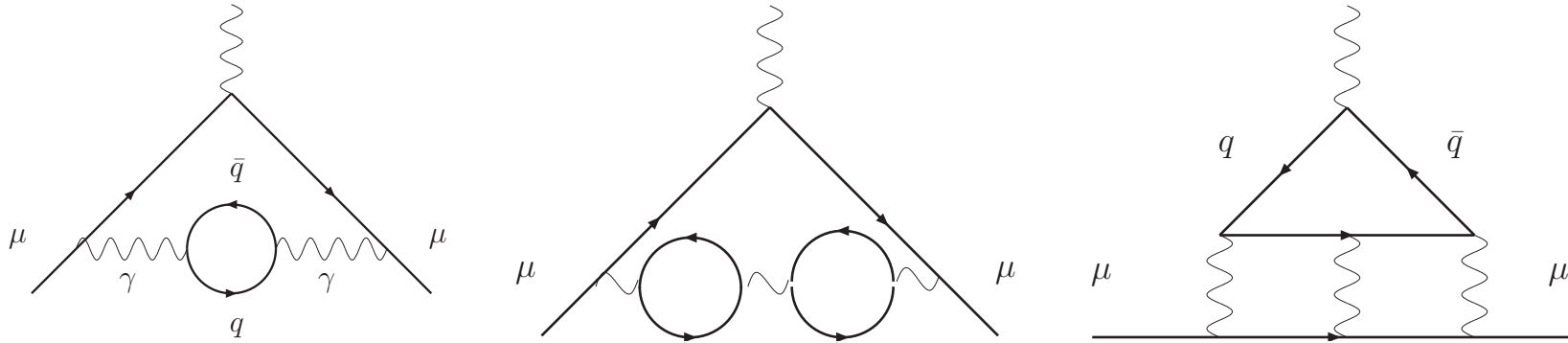
$$a_\mu^{QED} = 116584719(1) \times 10^{-11}.$$

This accounts up to four-loop contributions, i.e., up to the α^4 terms.

Kinoshita, Nio + Laporta, Remiddi

Hadronic contributions (1)

$$a_{\mu}^{had} = a_{\mu}^{had,LO} + a_{\mu}^{had,HO} + a_{\mu}^{had,LBL}$$



1. Lowest order hadronic contribution represented by a quark loop
2. An example of higher order hadronic contribution
3. Light-by-light scattering contribution

In fact

$$a_{\mu}^{had,LO} = \frac{1}{4\pi^3} \int_{4m_{\pi}^2}^{\infty} ds K(s) \sigma(s) = \frac{\alpha^3}{3\pi^2} \int_{4m_{\pi}^2}^{\infty} \frac{ds}{s} K(s) R(s),$$

where

$$K(s) = \int_0^1 dx \frac{x^2(1-x)}{x^2 + (1-x)s/m_{\mu}^2}$$

and

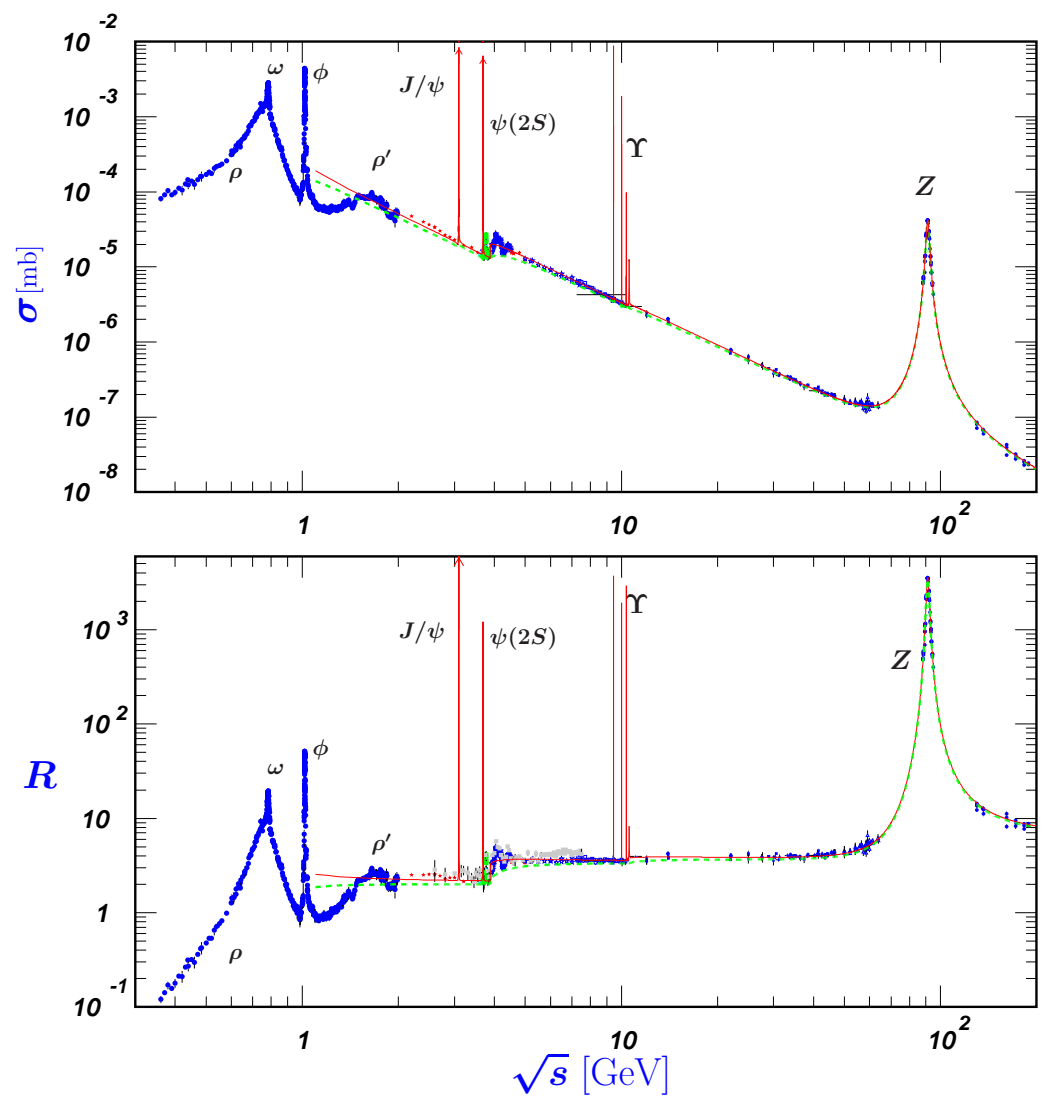
$$R(s) = \frac{\sigma(e^+e^- \rightarrow hadrons)}{\sigma(e^+e^- \rightarrow \mu^+\mu^-)}.$$

Close inspection shows that:

- $K(s) \sim 1/s$ at $s \rightarrow \infty$, so integral is saturated by low-energy region.
- About 91% of the total contribution to $a_{\mu}^{had,LO}$ is accumulated at center-of-mass energies $\sqrt{s} < 1.8 \text{ GeV}$.
- 73% of $a_{\mu}^{had,LO}$ is covered by the $\pi\pi$ final state, which is dominated by the $\rho(770)$ resonance.

Hadronic contributions (3)

Let's look at $R(s)$:



Hadronic contributions (4)

Hadronic LO contribution

$$\begin{aligned} a_{\mu}^{had,LO} &= 6909(44) \times 10^{-11} && \text{Davier; Eidelman 07} \\ &6894(46) \times 10^{-11} && \text{Hagiwara et al. 07} \\ &6921(56) \times 10^{-11} && \text{Jegerlehner 06} \\ &6944(49) \times 10^{-11} && \text{Troconiz, Yndurain 05} \end{aligned}$$

Hadronic HO contributions

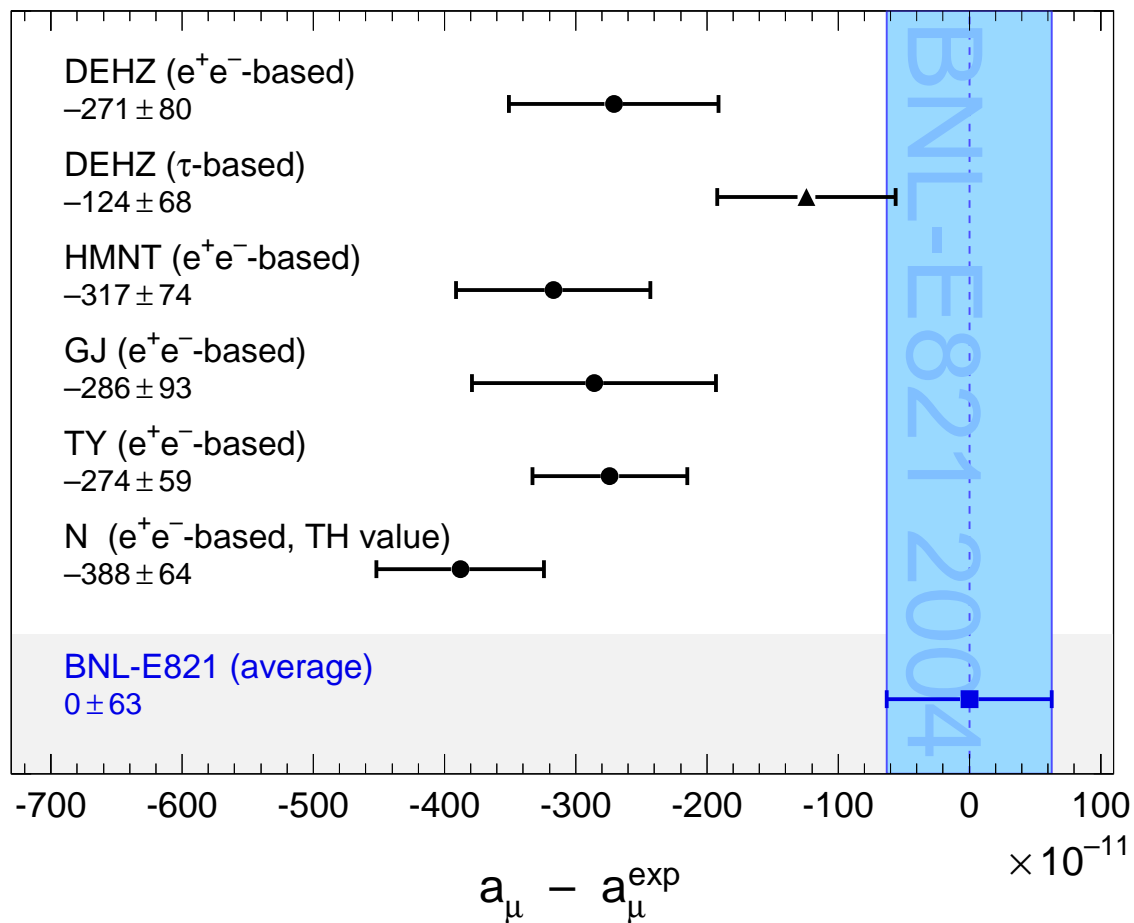
$$a_{\mu}^{had,HO} = -98(1) \times 10^{-11} \quad \text{Hagiwara et al. 07}$$

Electroweak contributions

$$a_{\mu}^{EW} = 154(2) \times 10^{-11} \quad \text{Czarnecki et al. 07}$$

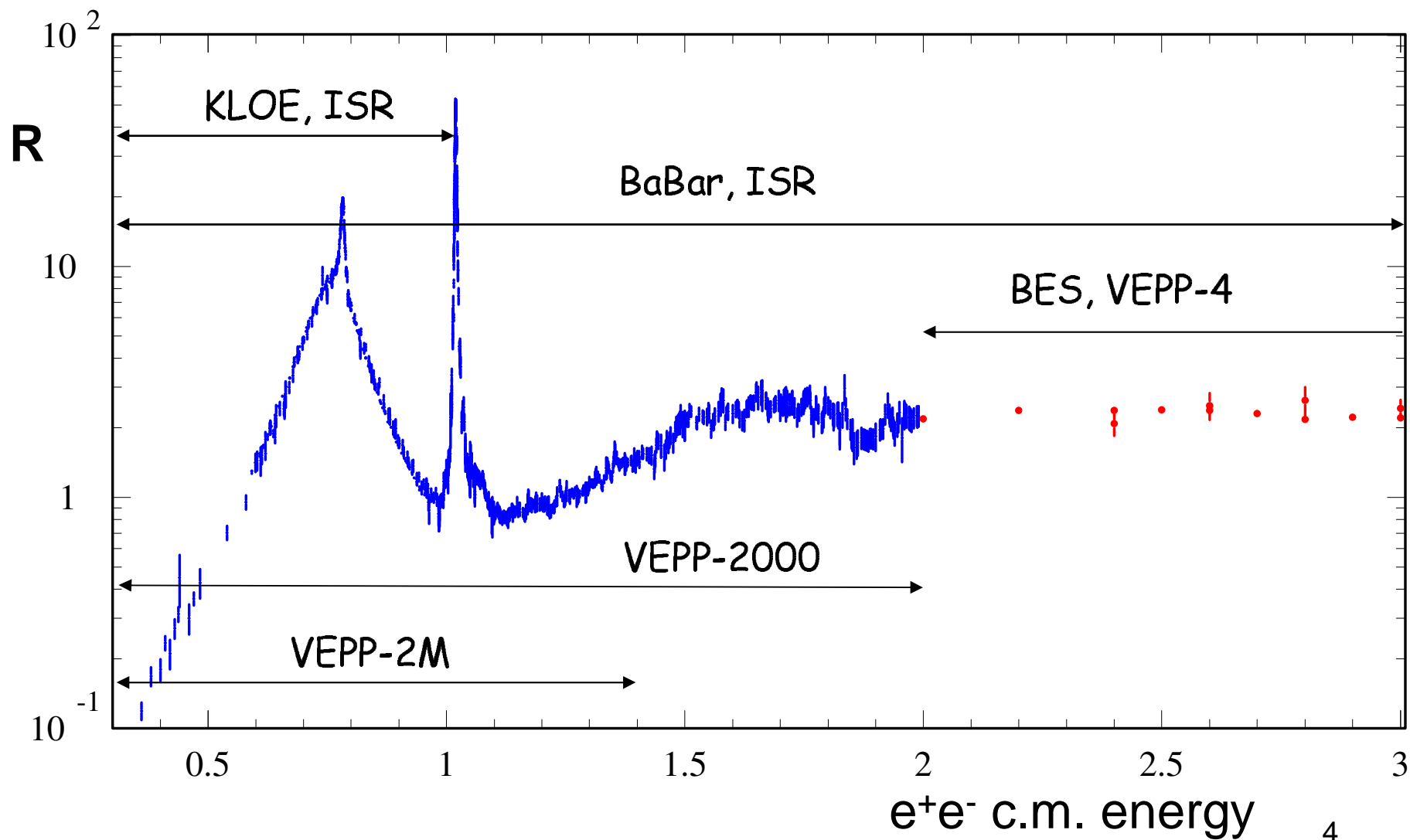
Light-by-light contribution needs some model considerations

$$\begin{aligned} a_{\mu}^{had,LBL} &= 80(40) \times 10^{-11} && \text{Knecht et al. 02} \\ &136(25) \times 10^{-11} && \text{Melnikov, Vainstein 04} \end{aligned}$$

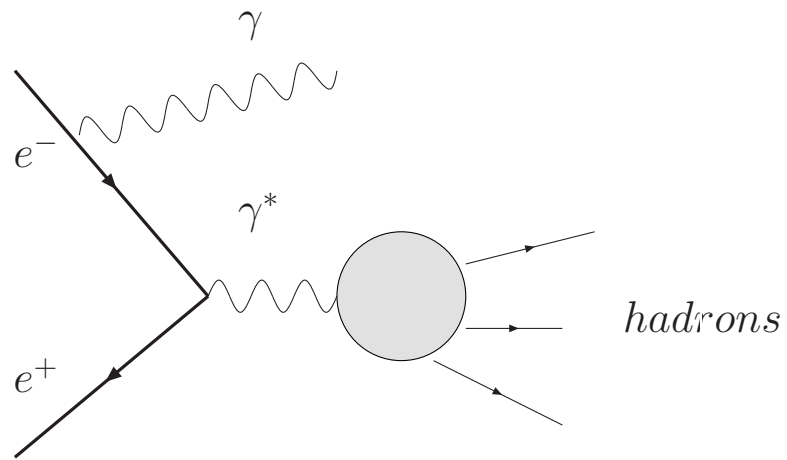


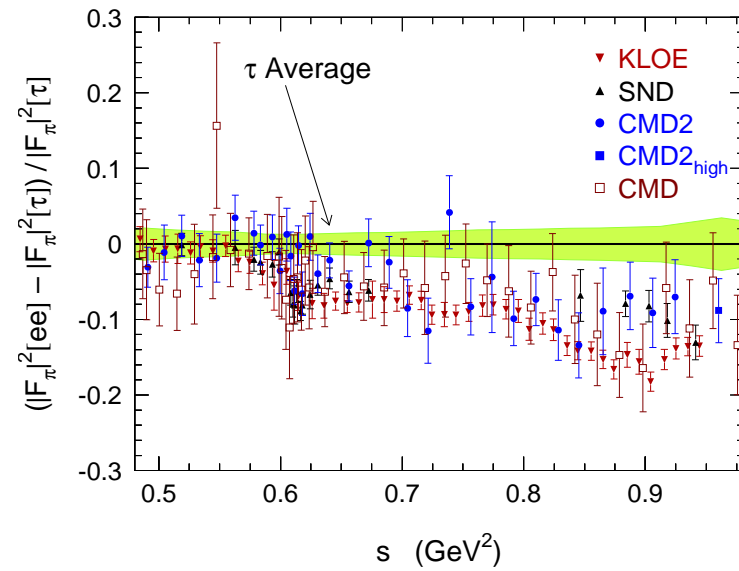
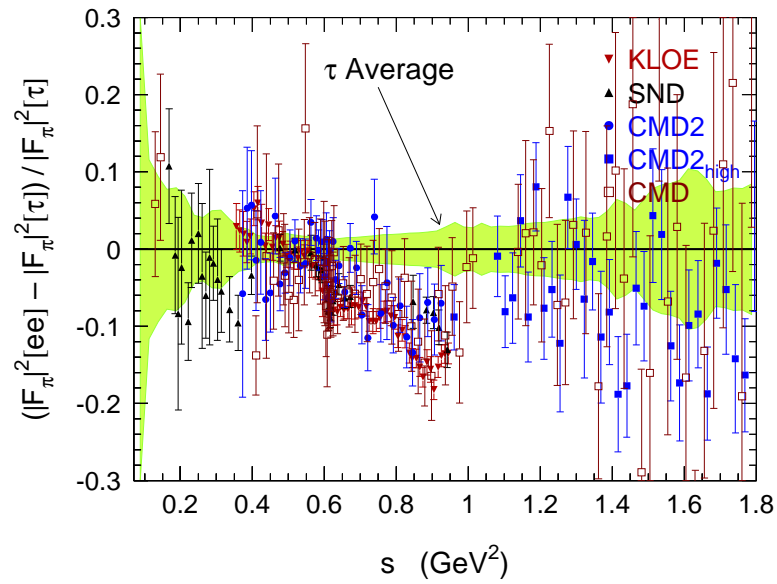
Typical value for difference:

$$a_\mu^{\text{exp}} - a_\mu^{\text{theor, SM}} = (280 \pm 80) \times 10^{-11} : \quad 3.4\sigma$$



ISR method





Oscillations in $D^0 - \bar{D}^0$ system

Time evolution in $D^0 - \bar{D}^0$ is described by Schrödinger equation

$$i \frac{\partial}{\partial t} \begin{pmatrix} D^0(t) \\ \bar{D}^0(t) \end{pmatrix} = \left(M - \frac{i}{2} \Gamma \right) \begin{pmatrix} D^0(t) \\ \bar{D}^0(t) \end{pmatrix},$$

where M and Γ matrices are Hermitian.

Mass eigenstates:

$$\begin{aligned} |D_1\rangle &= p |D^0\rangle + q |\bar{D}^0\rangle & |D_1(t)\rangle &= e^{-i(m_1 - i\Gamma_1/2)t} |D_1(0)\rangle \\ |D_2\rangle &= p |D^0\rangle - q |\bar{D}^0\rangle & |D_2(t)\rangle &= e^{-i(m_2 - i\Gamma_2/2)t} |D_2(0)\rangle \end{aligned}$$

Flavor states time evolution:

$$\begin{aligned} |D^0(t)\rangle &= e^{-(\Gamma_1/2 + im)t} \left\{ \cosh[(y + ix)\Gamma t/2] |D^0\rangle - \left(\frac{q}{p}\right) \sinh[(y + ix)\Gamma t/2] |\bar{D}^0\rangle \right\} \\ |\bar{D}^0(t)\rangle &= e^{-(\Gamma_1/2 + im)t} \left\{ \cosh[(y + ix)\Gamma t/2] |\bar{D}^0\rangle - \left(\frac{p}{q}\right) \sinh[(y + ix)\Gamma t/2] |D^0\rangle \right\} \end{aligned}$$

Oscillations in $D^0 - \bar{D}^0$ system

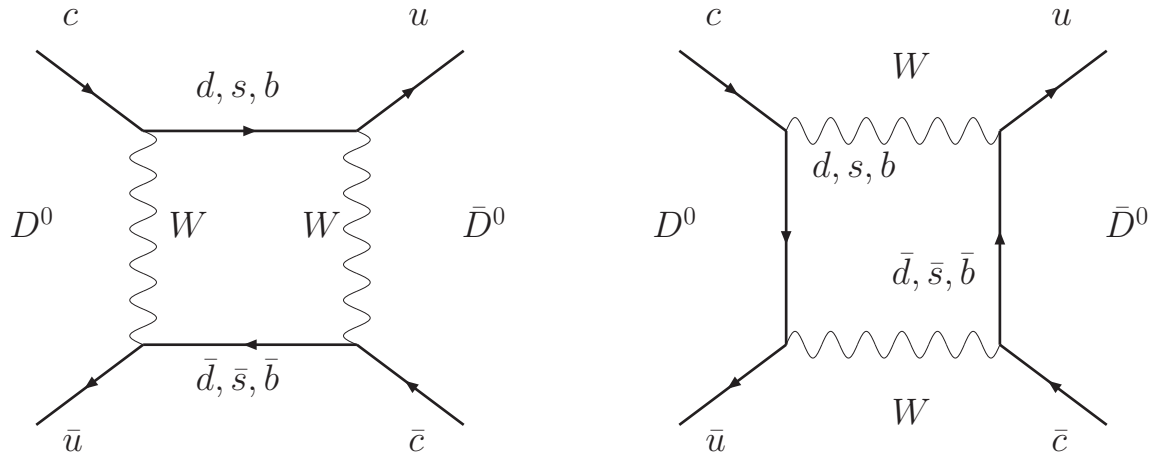
Here

$$m = \frac{1}{2}(m_1 + m_2), \quad \Gamma = \frac{1}{2}(\Gamma_1 + \Gamma_2),$$

$$x = 2 \frac{m_1 - m_2}{m_1 + m_2}, \quad y = \frac{\Gamma_1 - \Gamma_2}{\Gamma_1 + \Gamma_2},$$

Mixing exists if either x or y non-zero

Short-distance contributions in SM



is very small.

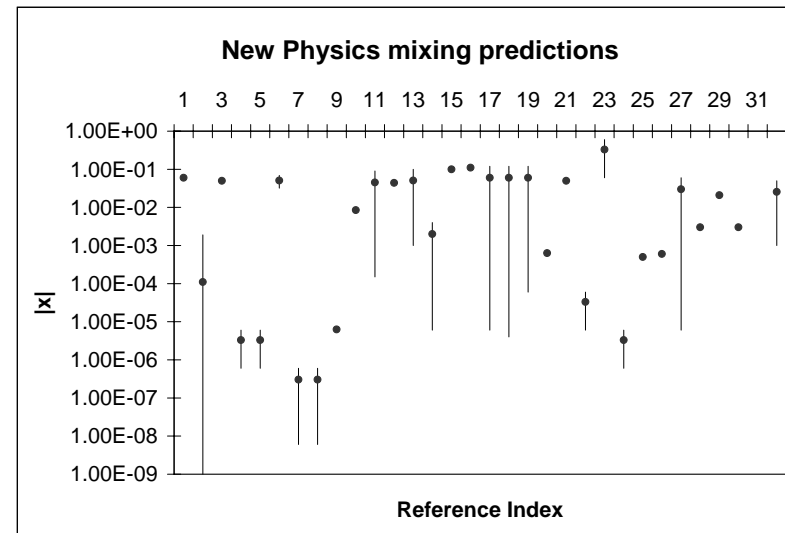
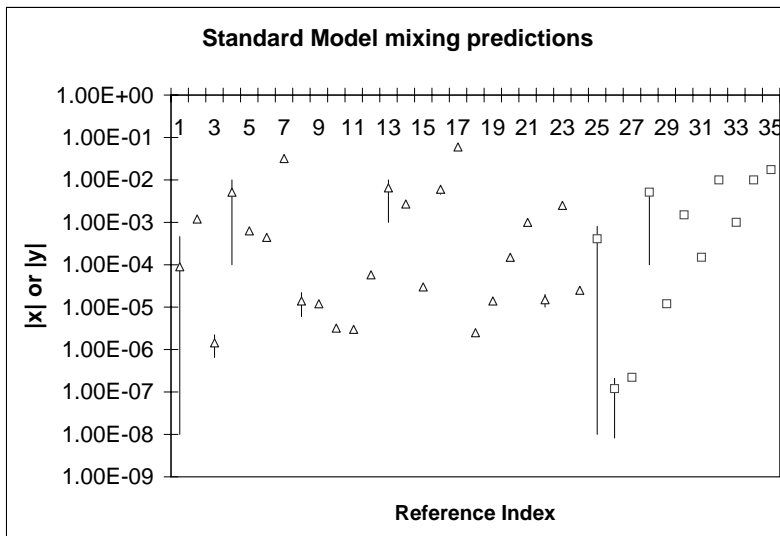
Oscillations in $D^0 - \bar{D}^0$ system

Reasons are:

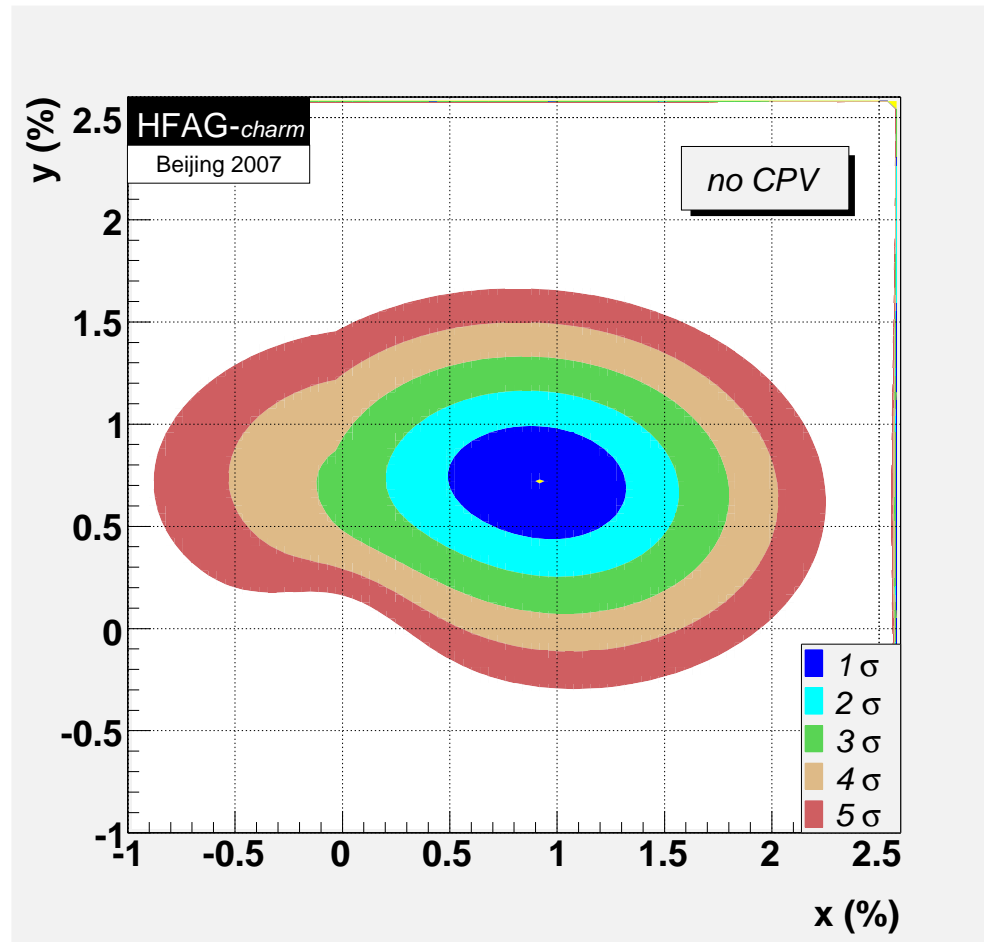
- b quark is CKM-suppressed
- s and d quarks are GIM suppressed
- $x \sim O(10^{-5})$ or less

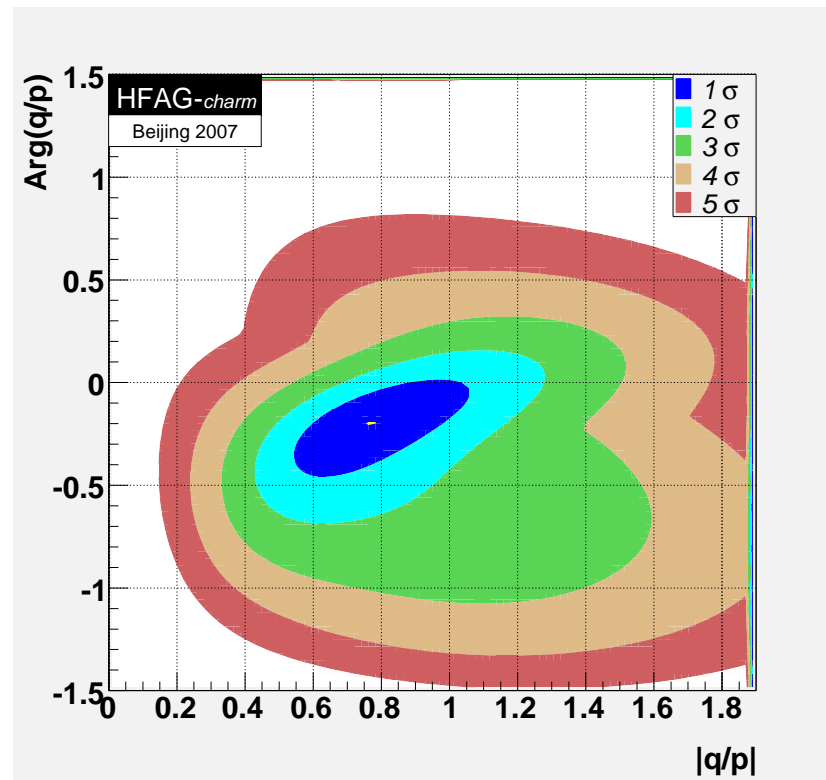
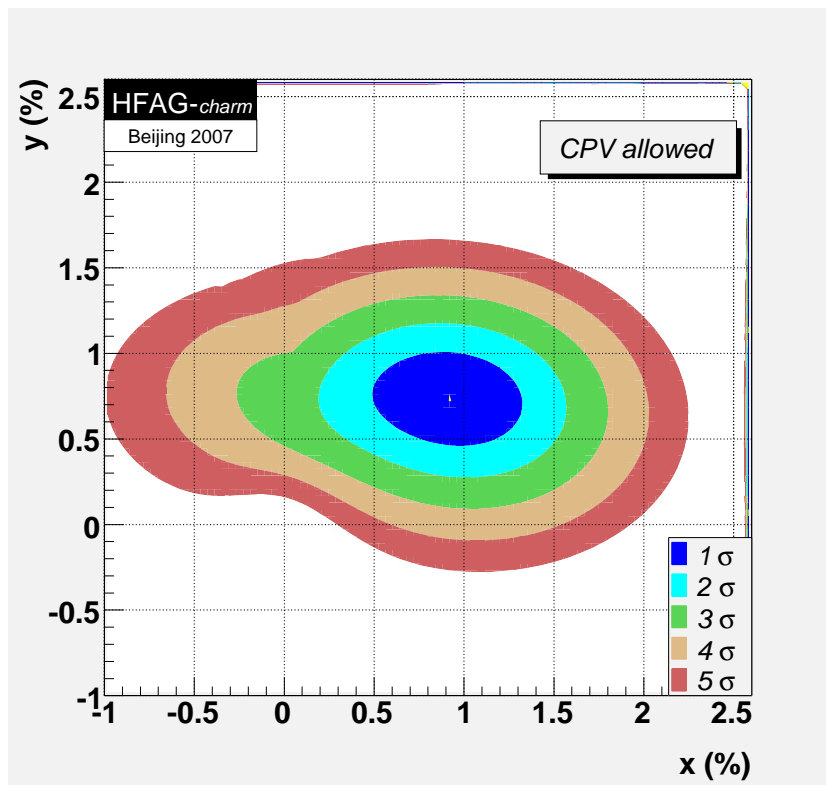
Long-distance contributions dominate but hard to estimate precisely.

Collection of predictions:



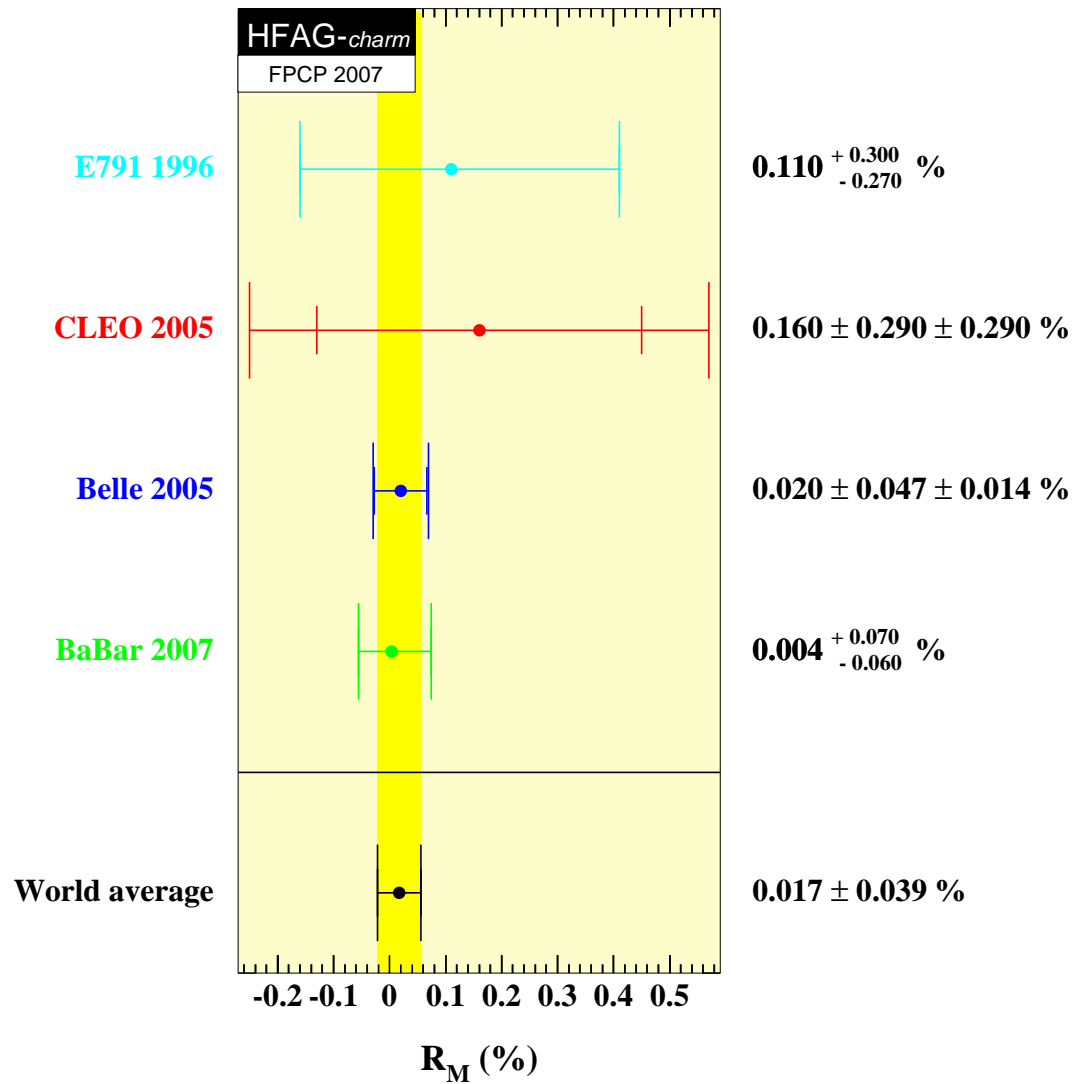
Heavy Flavor Averaging Group analyzed all world data. Most important are data of BELLE and BaBar.



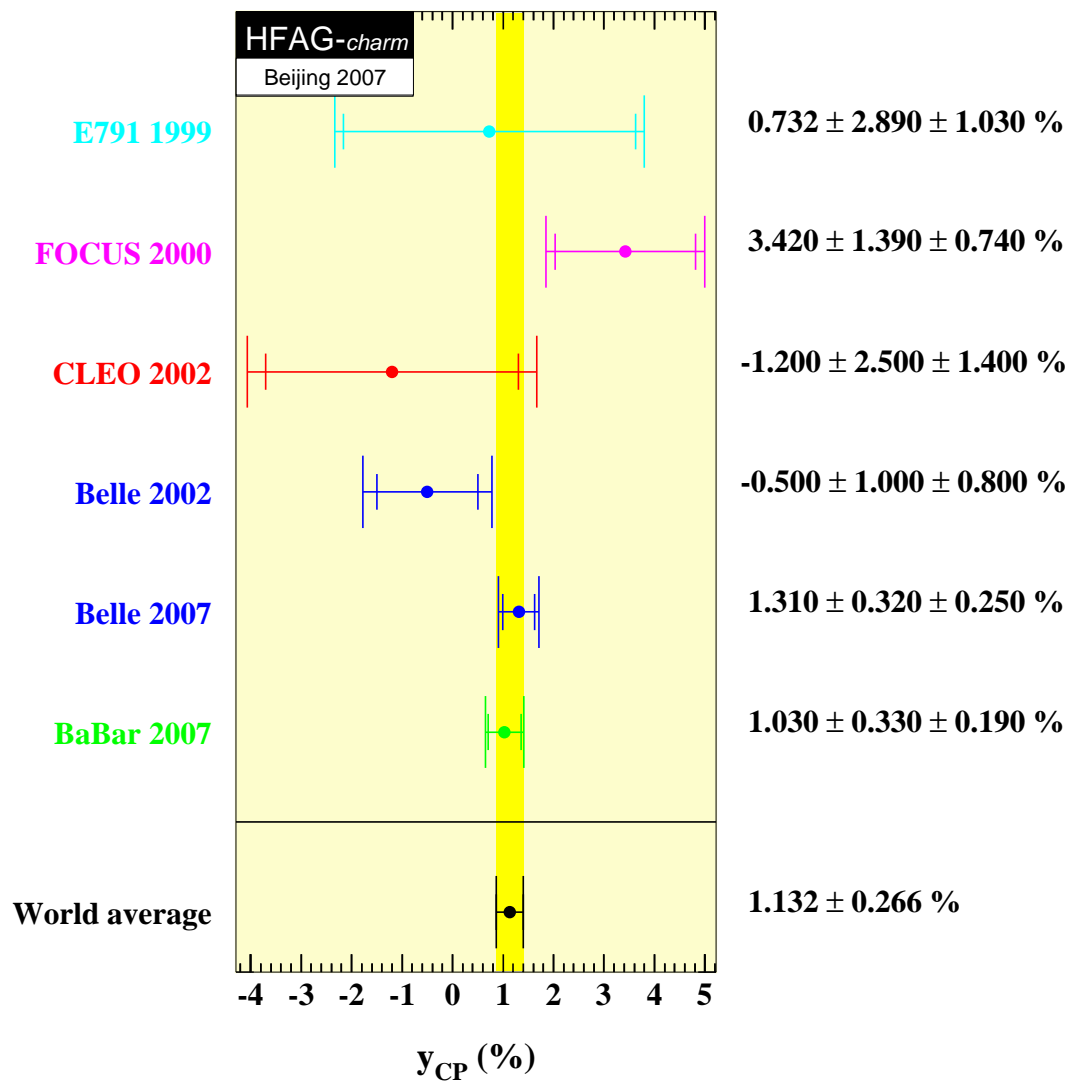


Conclusions of HFAG:

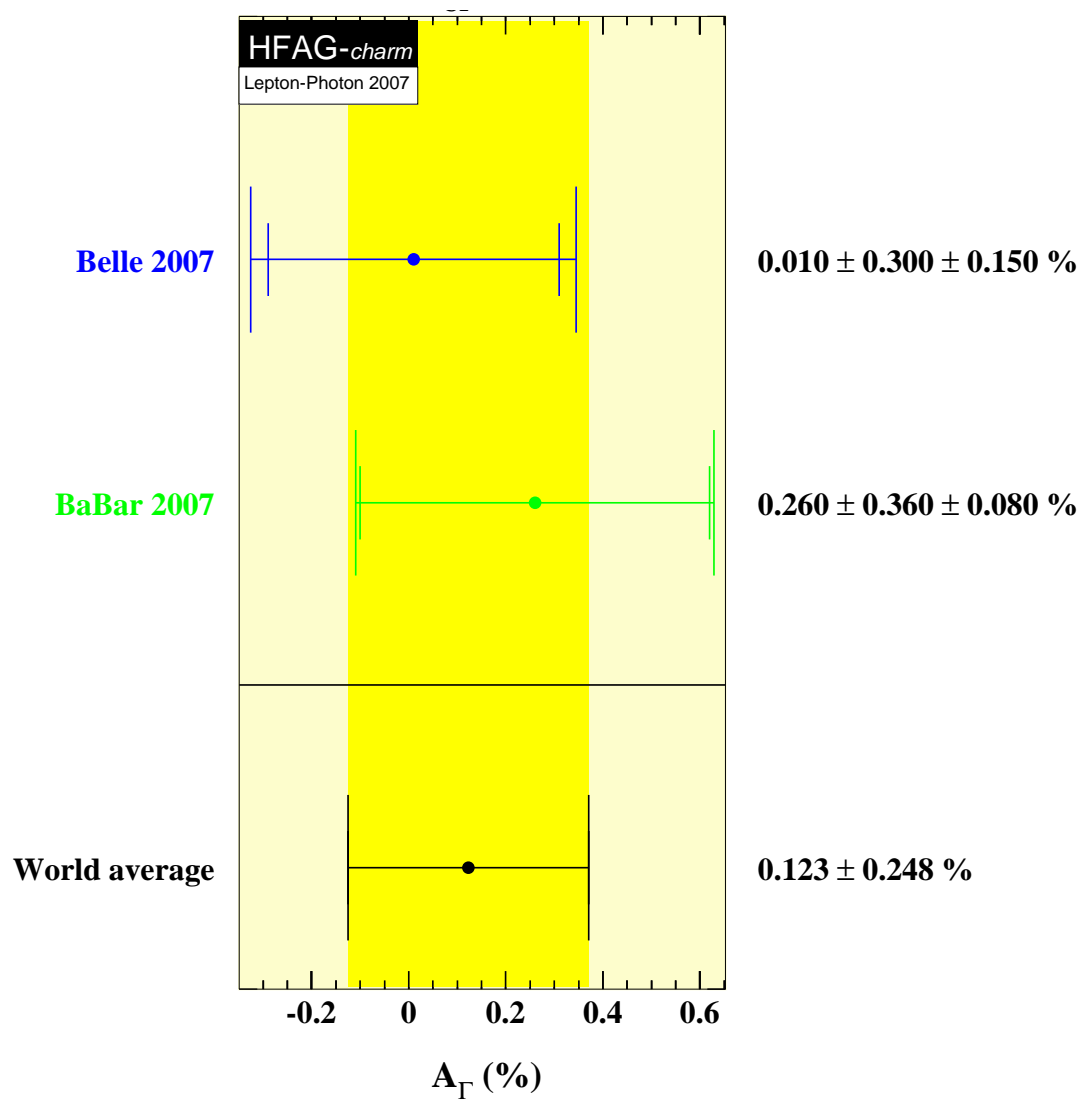
- The no-mixing point $x=0, y=0$ is excluded at 6.7σ . The effect is presumably dominated by long-distance processes, which are difficult to calculate.
- There is no evidence yet for CPV in the $D^0 - \bar{D}^0$ system. Observing CPV at the level of sensitivity of the current experiments would indicate new physics.



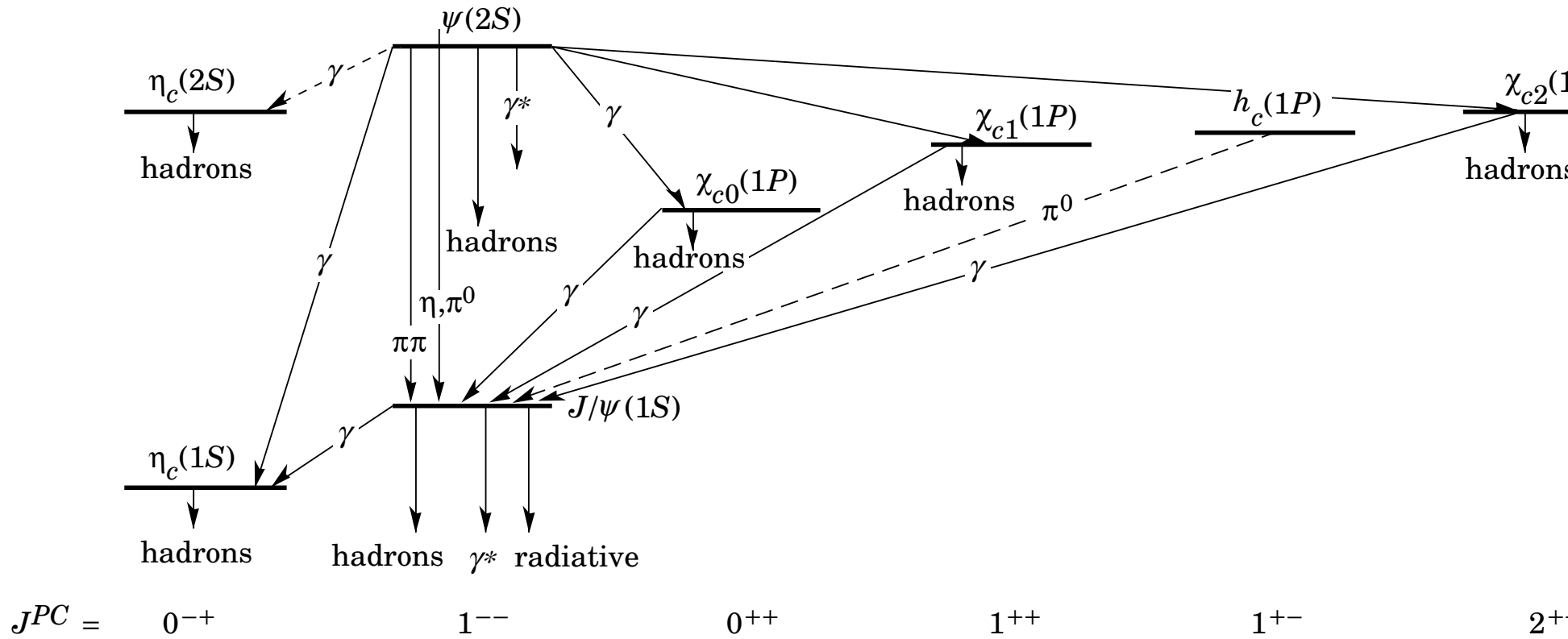
$$R_M = (x^2 + y^2)/2 \quad (3)$$



$$\tau_{CP} = \frac{\tau_{K\pi}}{\langle \tau_{hh} \rangle} - 1 \quad (4)$$



Charmonium



Bottomonium

